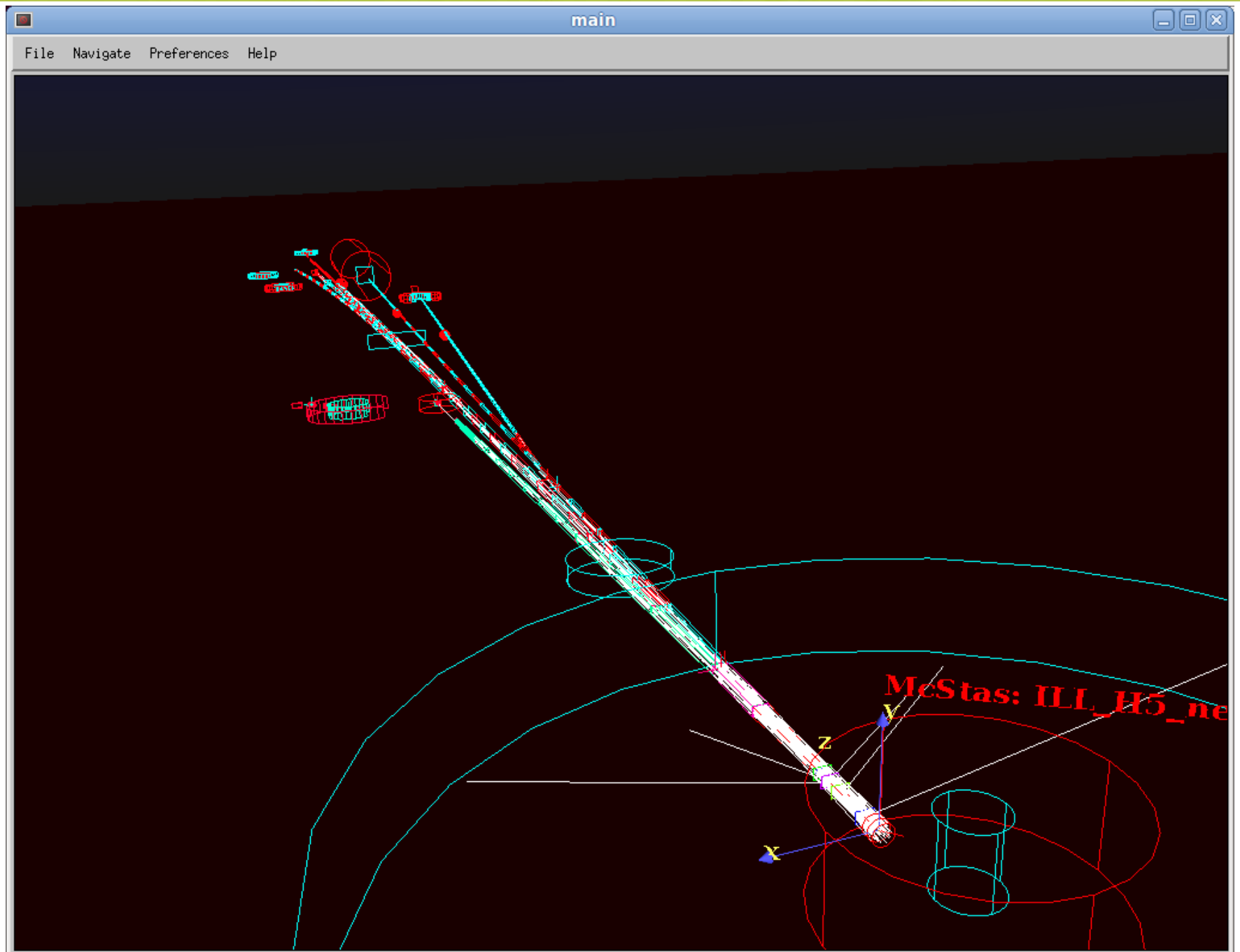


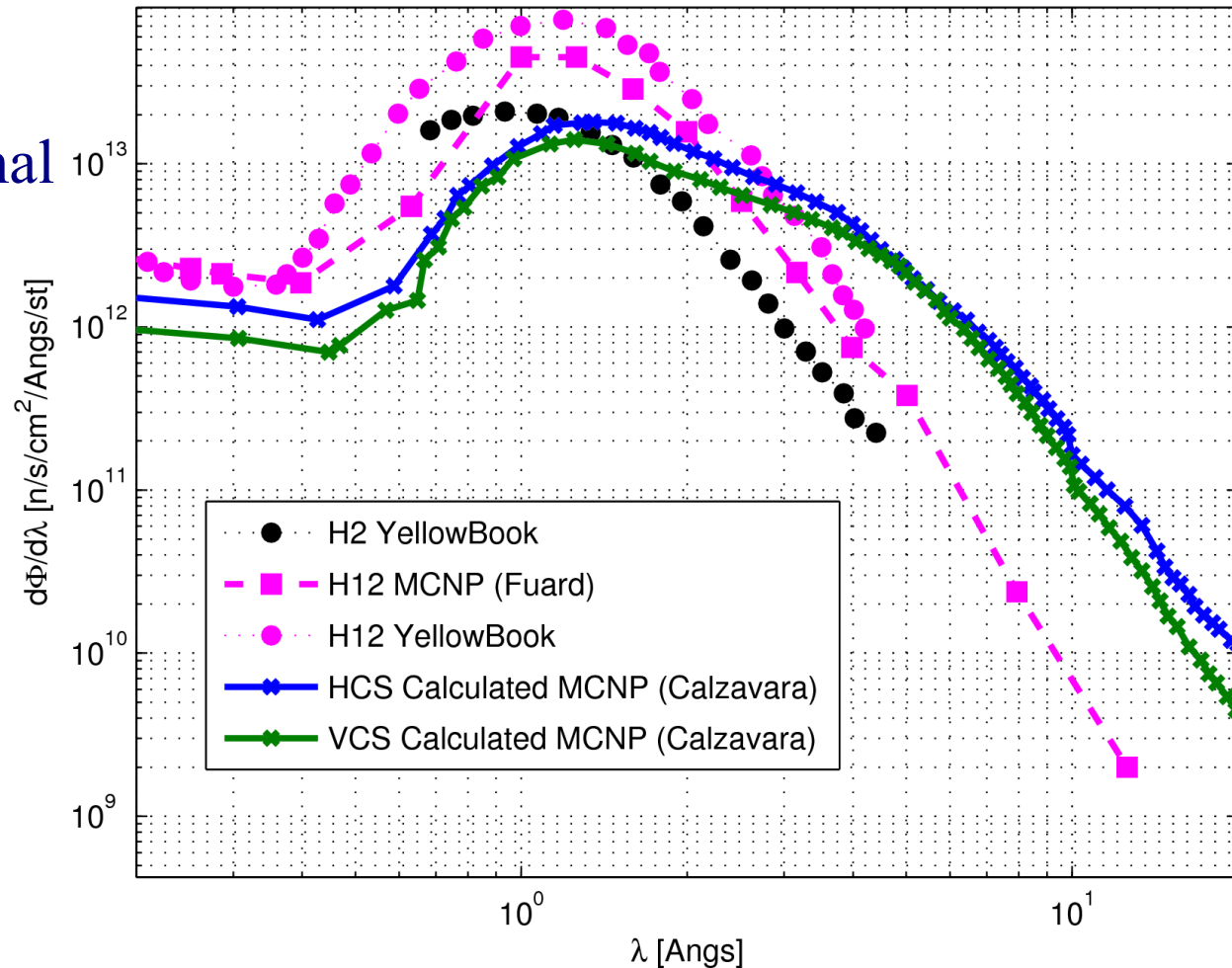
Advances on low energy moderator simulation using McStas: water, liq- ^4He , liq- D_2

E. Farhi, O. Zimmer, Y. Calzavara, ILL



We benchmarked the ILL thermal and cold sources.

Thermal $8 \cdot 10^{13}$ n/s/cm²/st/Å
 Cold: $2 \cdot 10^{13}$ n/s/cm²/st/Å

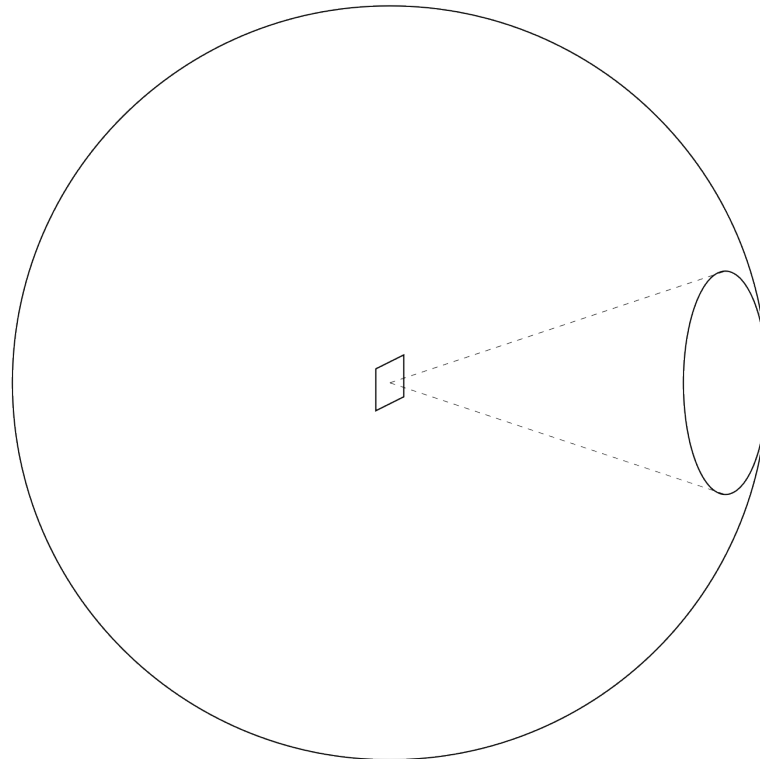


- ✚ Thermal source brilliance validated in agreement with the *YellowBook* data
- ✚ Cold sources are much **brighter** than reported in the *YB* by factor 3-10 vs λ
- ✚ Data and report published at <http://www.ill.eu/?id=11169>
- ✚ FRM2 and HZB data also available

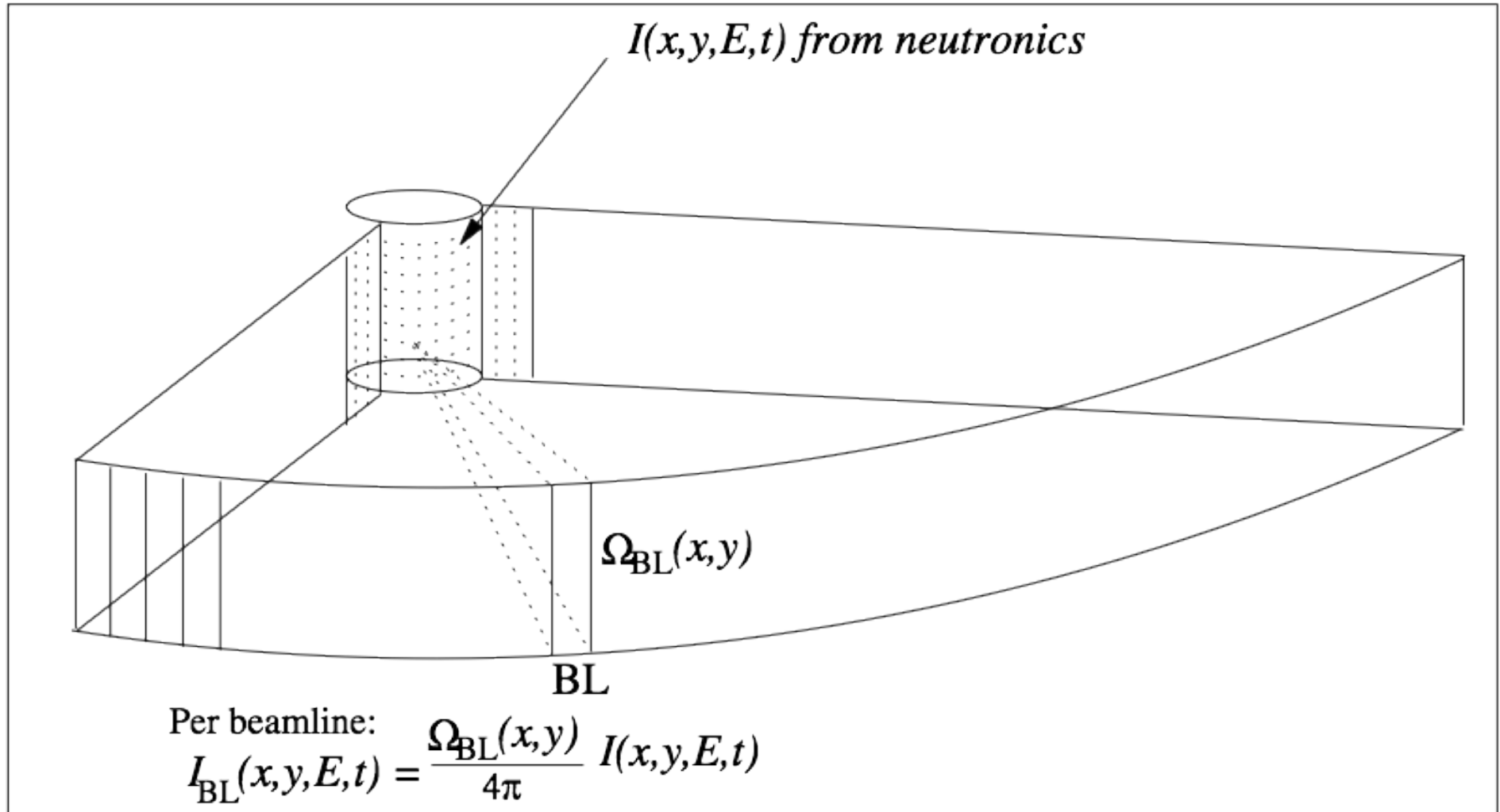
Source brilliance

- Reasonable unit to compare emittance of sources:

$$\mathfrak{B} [n/cm^2/s/ster/\text{\AA}]$$



Moderators... (Where McStas starts)



The Brilliance_monitor Component

Special "Brilliance" monitor of FIXED size 1x1cm. If used in the right setting, will output "instantaneous" and "mean" brilliances in units of Neutrons/cm²/ster/AA/s. Conditions for proper units:

- Use a with a source of area 1x1cm
- The source must illuminate/focus to an area of 1x1cm a 1m distance
- Parametrise the Brilliance_monitor with the frequency of the source
- To not change the source TOF distribution, place the Brilliance monitor close to the source!

with a source of area 1x1cm illuminating/focusing to an area of 1x1cm a 1m distance, this monitor will output "instantaneous" and "mean" brilliances in units of Neutrons/cm²/ster/AA/s

Identification

- **Author:** Peter Willendrup, derived from TOF_lambda_monitor.comp
- **Origin:** DTU Physics
- **Date:** May 23, 2012
- **Version:** 1.1

Description

Here is an example of the use of the component. Note how the mentioned Unit conditions are implemented in instrument code.

```
COMPONENT Source = ESS_moderator_long(
  l_low = lambdamin, l_high = lambdamax, dist = 1, xw = 0.01, yh = 0.01,
  freq = 14, T=50, tau=287e-6, tau1=0, tau2=20e-6,
  n=20, n2=5, d=0.00286, chi2=0.9, I0=6.9e11, I2=27.6e10,
  branch1=0, branch2=0.5, twopulses=0, size=0.01)
AT (0, 0, 0) RELATIVE Origin

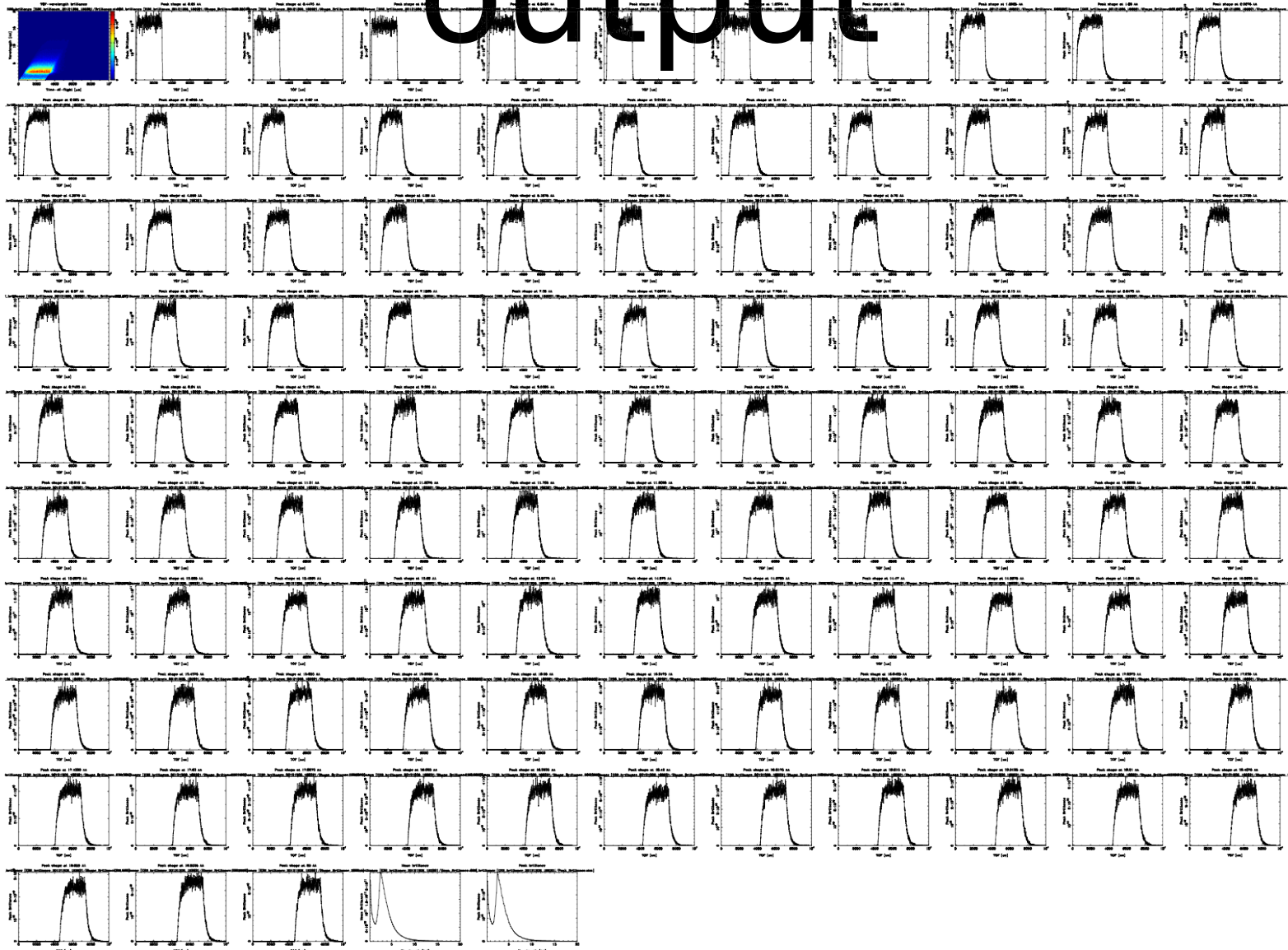
COMPONENT BRIL = Brilliance_monitor(nlam=196,nt=401,filename="bril.sim",
  t_0=0,t_1=4000,lambda_0=lambdamin,
  lambda_1=lambdamax, Freq=14)
AT (0,0,0.000001) RELATIVE Source
```

Input parameters

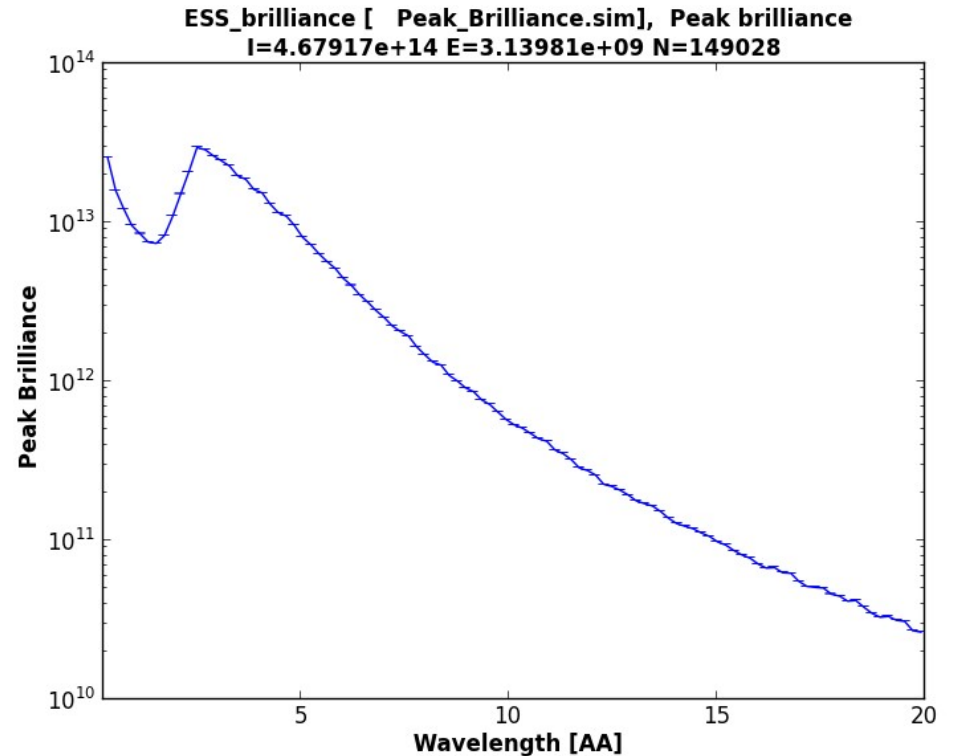
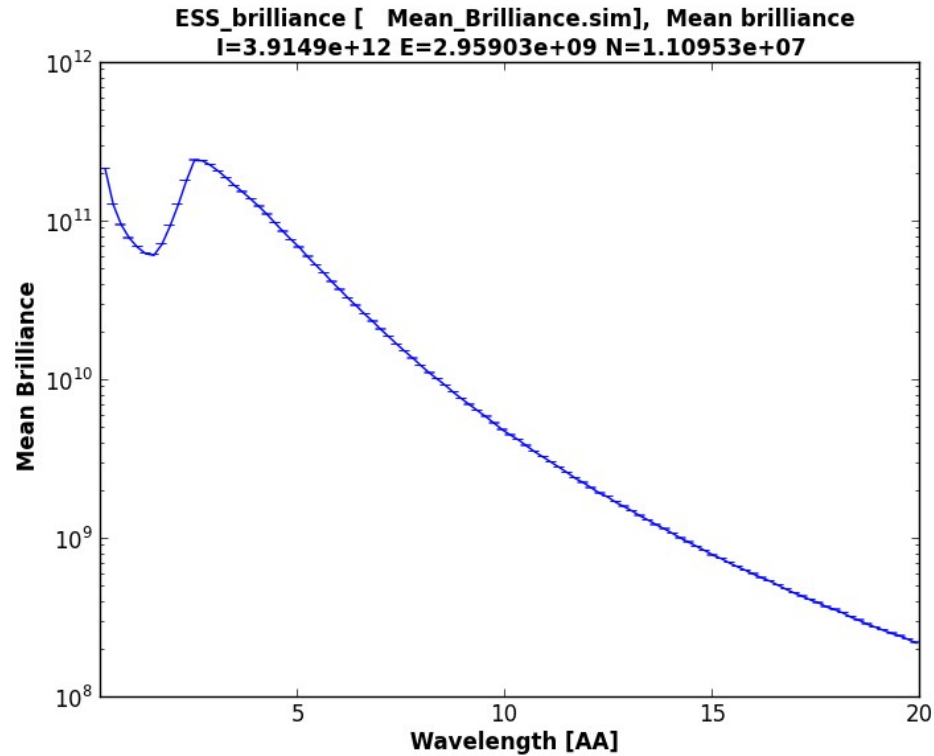
Parameters in **boldface** are required; the others are optional.

Name	Unit	Description	Default
nlam	1	Number of bins in wavelength	101
nt	1	Number of bins in TOF	1001
filename	string	Defines filenames for the detector images. Stored as: Peak_<filename> and Mean_<filename>	
t_0	us	Minimum time	0
t_1	us	Maximum time	20000
lambda_0	AA	Minimum wavelength detected	0
lambda_1	AA	Maximum wavelength detected	20
restore_neutron	1	If set, the monitor does not influence the neutron state	0
Freq	Hz	Source frequency. Use freq=1 for reactor source	

Brilliance_monitor output



Easy way to get brilliance curves



The moderation process implies a series of slow-down processes:

- Elastic collision with large momentum transfer – Compton scattering ($q \rightarrow \text{Inf}$, high energy neutrons)
- Elastic scattering with lower momentum transfers ($q \leq \text{few } 10 \text{ \AA}^{-1}$)
- Inelastic scattering on material dynamics ($E_i \leq \text{few } 100 \text{ meV}$)

Usual materials used for slowing down neutrons:

- Hydrogenated materials – water to get thermal neutrons (pre-moderator)
- Mesitylene (10-100K) – to get warm neutrons
- Liquid D₂ and H₂ – to get cold neutrons, using the rotational lines (e.g. 7-14meV)
- Solid D₂ – same as above but heat extraction is less efficient (no flow)
- Liquid ⁴He – to get ultra cold neutrons from $\sim 9 \text{ \AA}$ neutrons (post-moderator)

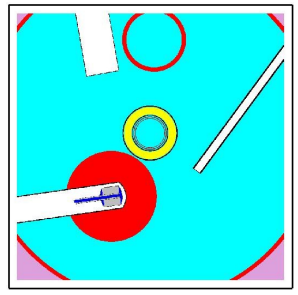
Instrument: Source: Generating neutron events

Neutron events may be currently generated from:

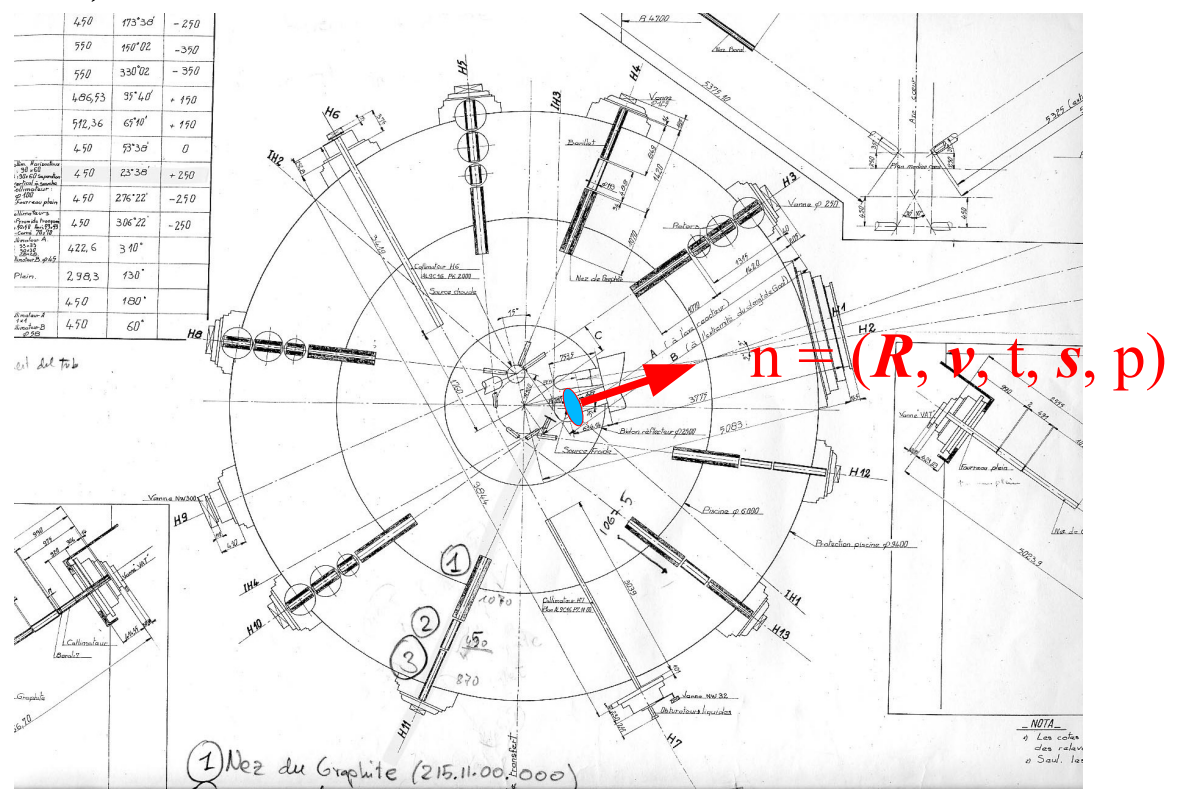
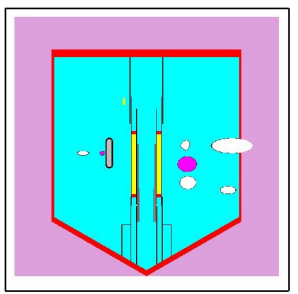
- A black body (Maxwellian distributions)
- Measured flux/divergence distributions
- MCNP and Tripoli (nuclear reactor simulation codes) event files

Other neutron generators are in principle easy to write (e.g. from GEANT4, FLUKA, ...)

top



side



So what's wrong with it ?

Available material neutron scattering cross sections for in ENDF:

liq-D2, liq-H2, H2O, D2O, liq-CH4, sol-CH4
ZrH_x, Graphite, Be, BeO
(that's all folks)

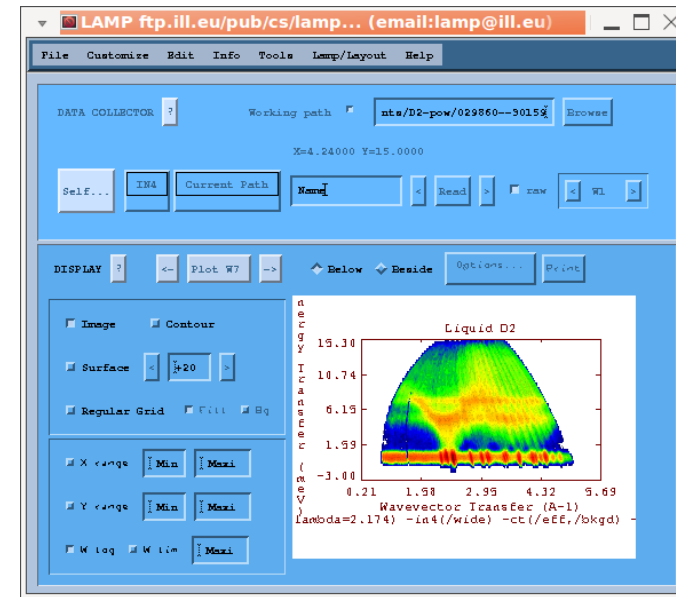
Weaknesses:

- Hardly produce Bragg peaks.
 - Rely on analytical models at low energies.
 - Mostly assumed to be incoherent-like.
 - Very few temperatures available for each material.
 - Data files often given for few discrete scattering angles (!!)
- this is ok when averaging on many scattering events in a large volume, but not accurate.

Improving these aspects also benefits to the neutron scattering community by setting **new demand on material simulation accuracy** in our MC codes.

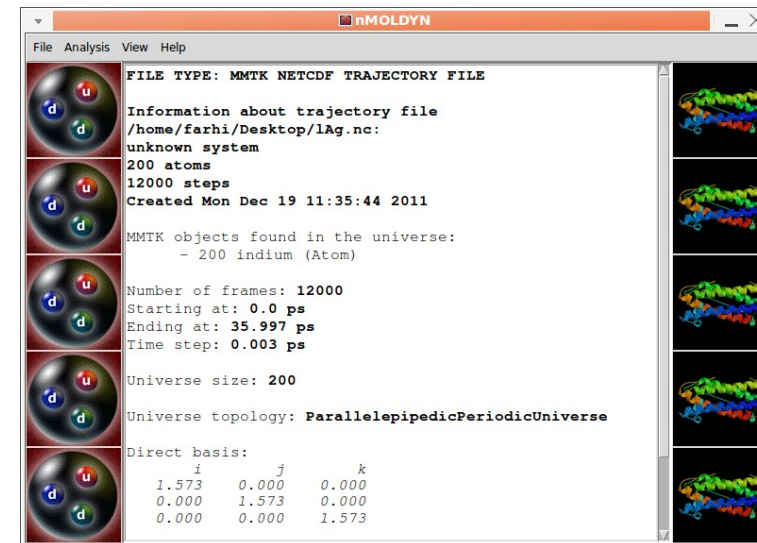
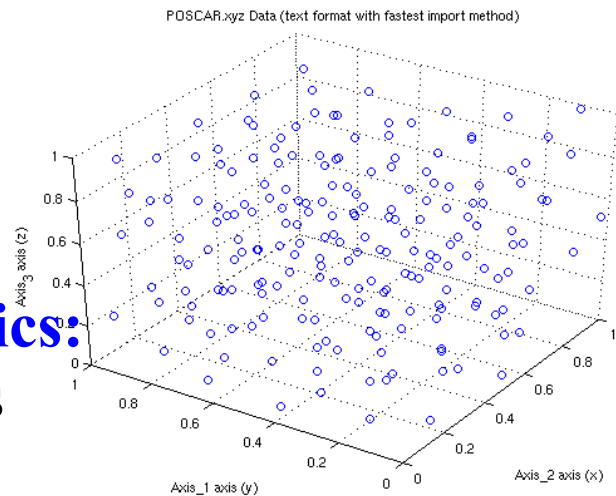
From experiments:

- Data mine or send new proposals
 - Reduce data: normalize, subtract empty cell,
 - Transform to $S(q, w)$
- I use **LAMP** for this.



From molecular dynamics:

- Perform MD simulations
- Import trajectories
- Compute coherent and incoherent scattering $S(q, w)$



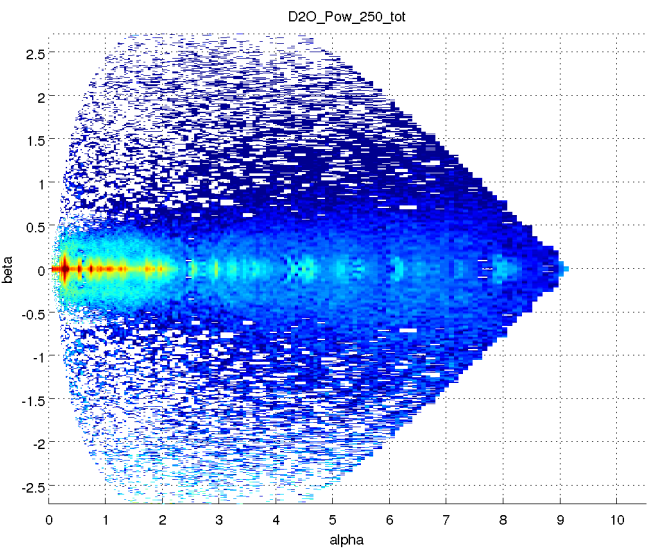
Mostly for high temperature materials – low temperatures are often quantum...
 I use **VASP**, then **nMoldyn** for this. May use Material Studio.

Measurement on IN4 Farhi/Calzavara/Haeck

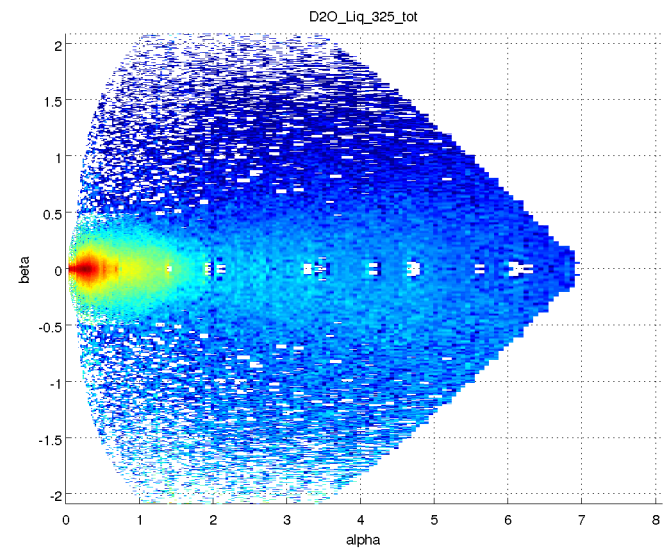
Exp 1-10-9: IN5 13-15 Sept 2010: H2O and D2O T=2-300 K; lambda=2,5,10 Angs

Use LAMP to generate $S(q,w)$. Use iFit to convert to $S(a,b)$.
 Use NJOY to convert to ACE.

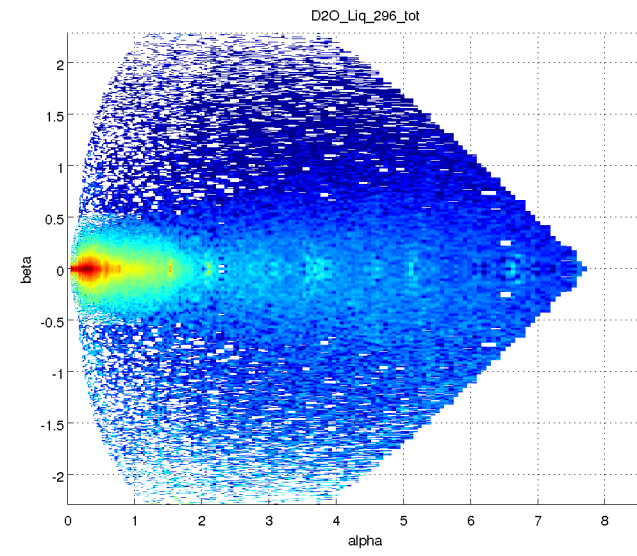
$$\alpha = \frac{E'+E - 2\sqrt{E'E} \cos\theta}{AkT} = \frac{\hbar^2 \kappa^2}{2M kT} \quad \beta = \frac{E'-E}{kT} = \frac{\varepsilon}{kT} \quad S(\alpha, \beta) \sim S(q, \omega) \cdot q \cdot M \cdot (k_B T)^2$$



250K powder

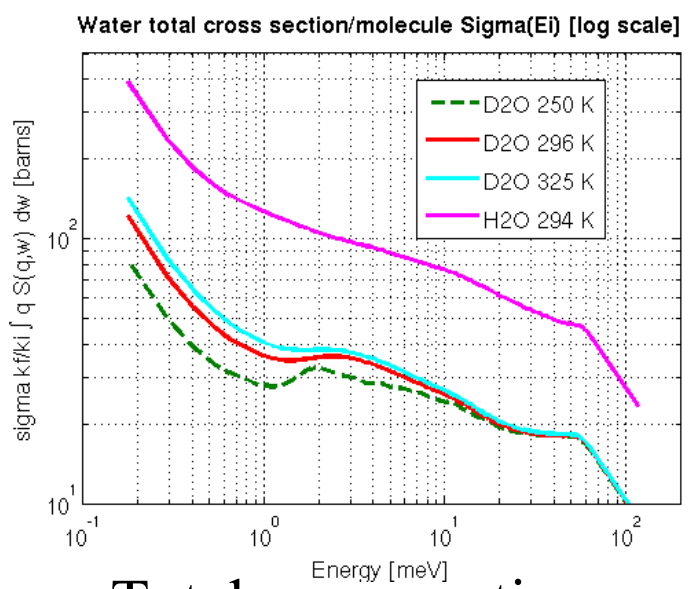


296K liquid

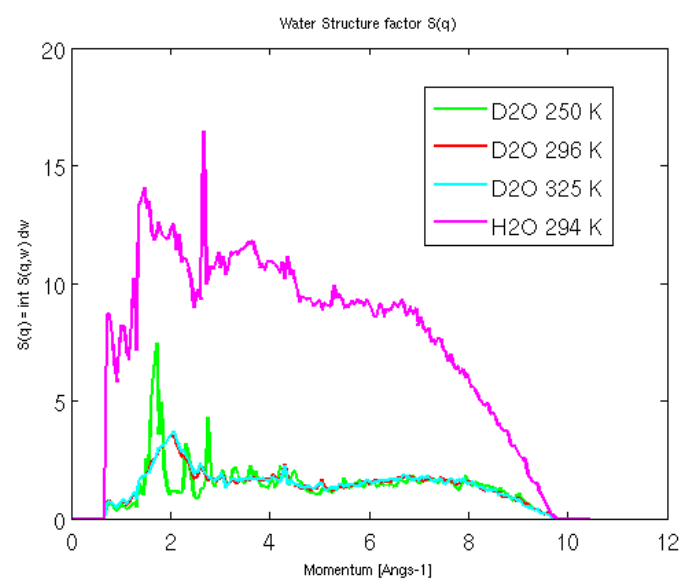


325K liquid

Water: comparing with existing data



Total cross section



Structure factor

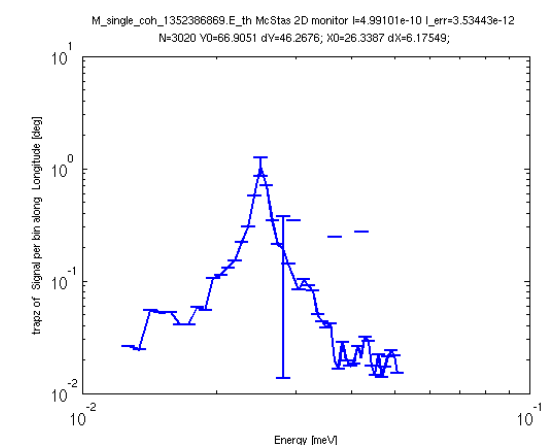
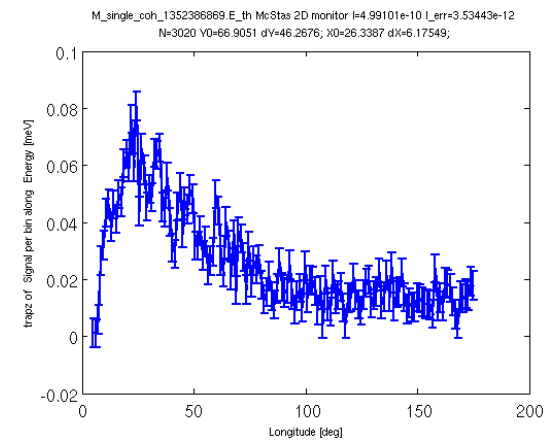
We have good data. Will definitely improve data bases and moderator models.



INTERNATIONAL ATOMIC ENERGY AGENCY
INDC(NDS)-0470
Distr. AC

INDC INTERNATIONAL NUCLEAR DATA COMMITTEE

Thermal Neutron Scattering Data
for the Moderator Materials



Tools used for this study

Upgraded Isotropic_Sqw component:

- More accurate total cross section computation $\sigma(Ei)$
- Computes many physical parameters from $S(q,w)$
- Compressibility, sound velocity, $\langle u^2 \rangle$, elastic modulus, diffusion coefficient, collective excitation frequency w^2 , gDOS, ...
- Can handle both classical and quantum limit $S(q,w)$ and symmetrize energy range, with Bose factor

iFit for handling the data

- Transformation $S(q,w) \rightarrow S(\alpha,\beta)$
- Plotting

See Egelstaff or H. Fischer, *Rev. Prog. Phys.* **69** (2006) 233

Neutronist's Mantra

$$\frac{d^2\sigma}{d\Omega dE_f} = \frac{\sigma}{4\pi} \frac{k_f}{k_i} N S(q, \omega)$$

Holy Book (Squires)

Effective cross section
for scattered intensity

$$\hat{\sigma} = \iint \frac{d^2\sigma}{d\Omega dE_f} d\Omega dE_f$$

V.F. Sears. *Adv. Phys.*, 24, 1, 1975.

We like to play games
in (q, ω) space

$$\frac{d\Omega}{d\theta} = -2\pi \sin\theta$$

$$\frac{dq}{d\theta} = -\frac{k_i k_f \sin\theta}{q}$$

Effective cross section
in (q, ω) space

$$\hat{\sigma} = \sigma \iint \frac{S(q, \omega) q}{2k_i^2} dq d\omega$$

Probability to interact

$$p = e^{-\rho \hat{\sigma} x}$$

Scattering distribution

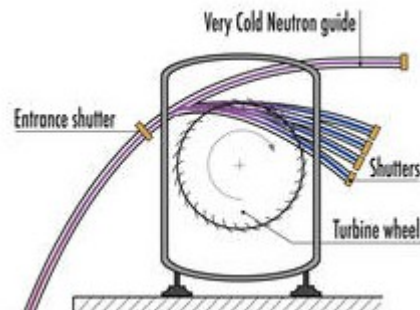
$S(q, \omega)$

with importance sampling to
scatter preferably where S is large

Cold neutrons are generated from moderation in e.g. I -H or I -D.

Ultra Cold Neutrons, produced from cold neutrons, used for fundamental physics (cosmology). Require high UCN density.

Currently max density at the ILL 30 cm^{-3} using a turbine (Doppler effect).

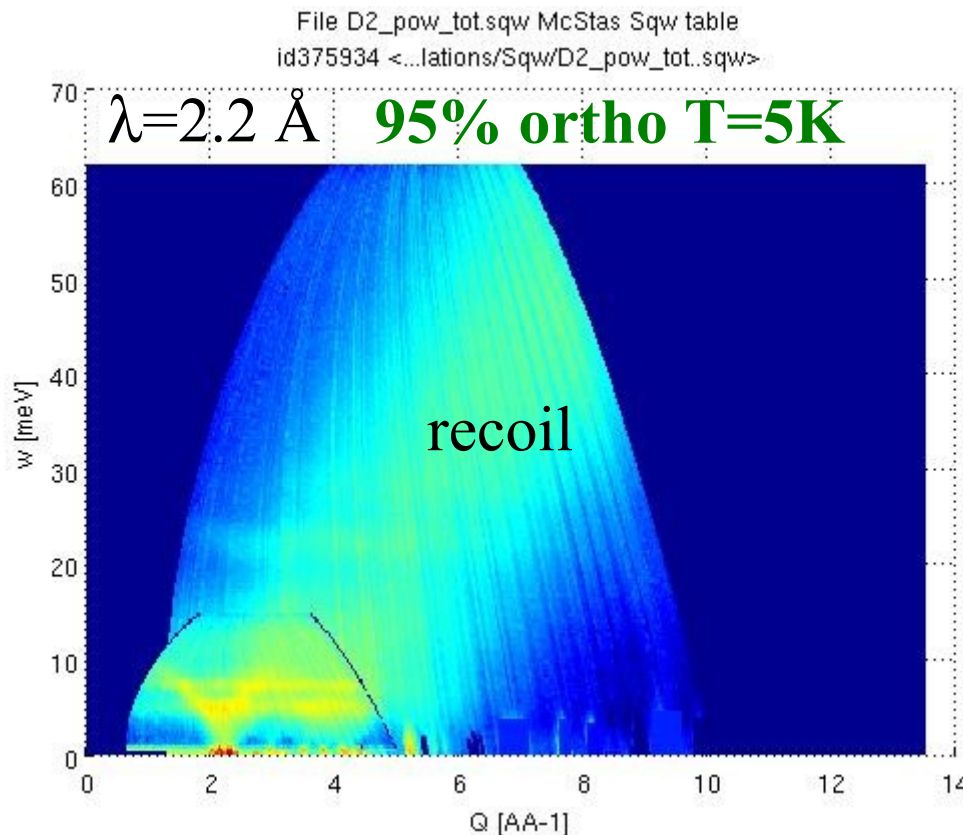


Idea : extract energy from cold neutrons using phonons at low temperature. Neutron can not regain energy because T is very low (\ll phonon energy).

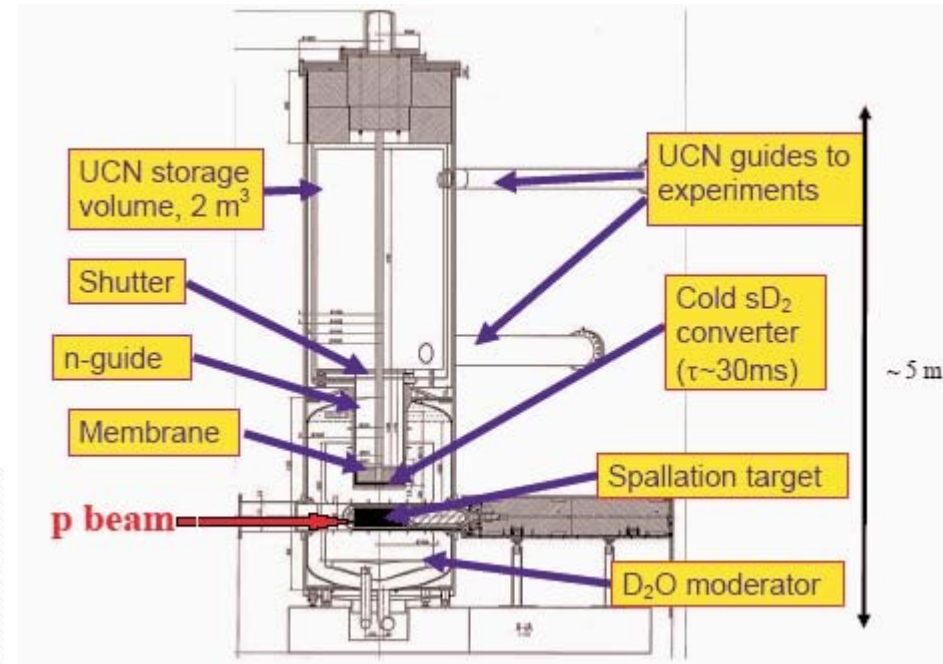
Materials : solid-D₂, ⁴He are promising

solid Deuterium as converter (o-D₂):

- temperature: 5 K
- para-D₂ content < 1%
- absorption cross section $\sigma_{\text{abs}} = 0.0005\text{b}$
- incoherent (2b) and coh (5b).
- Distribution of inelastic processes.



Used e.g. at PSI for the **nEDM**.



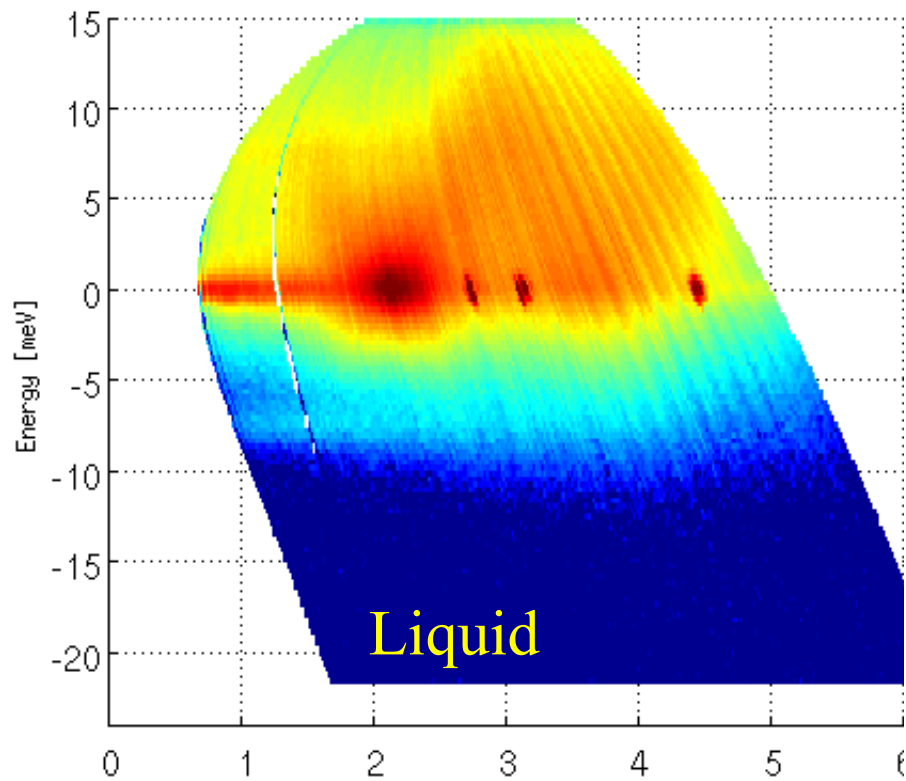
Data from Gutschmiedl (2007).

1 hour@IN4

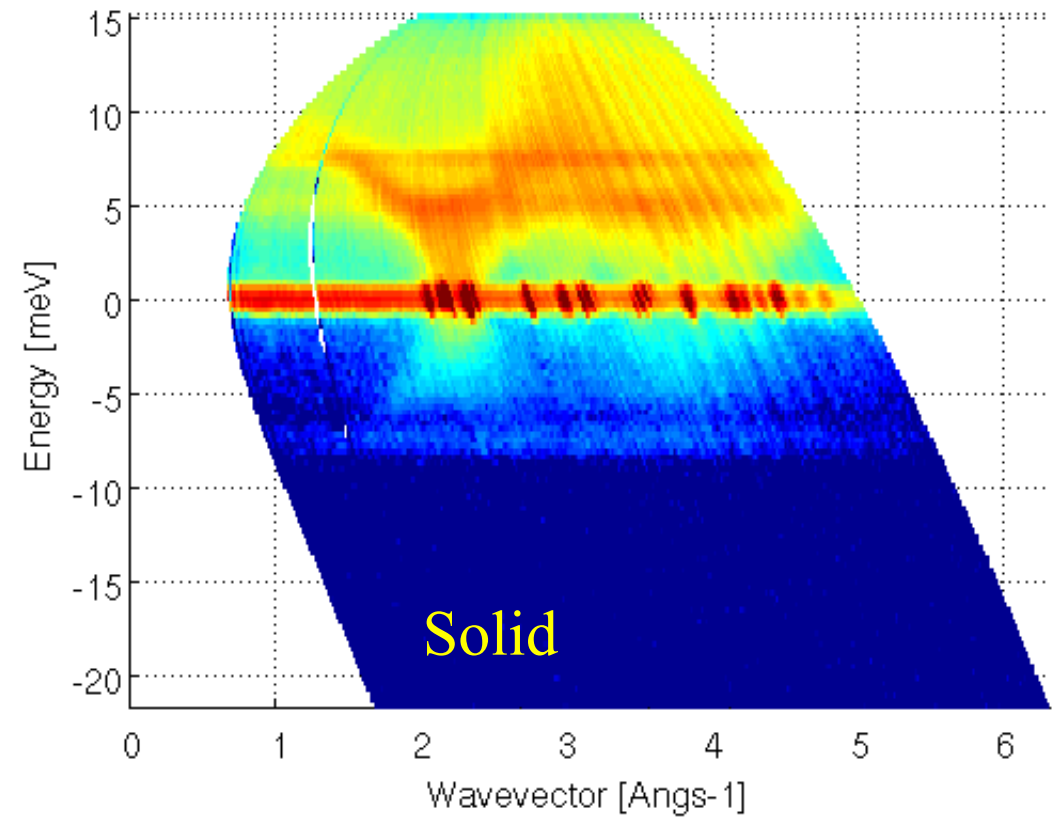
- A. Frei, E. Gutschmiedl, et al. *EPL*, **92** (2010) 62001
- Granada, *EPL* **86** (2009) 66007

May be converted to $S(\alpha, \beta)$
to complement MCNP kernels
Also for $l\text{-D}_2$

liq-D₂: 20K IN4@ILL Gutmiedl Nov 2007



sol-D₂: 5K IN4@ILL Gutmiedl Nov 2007



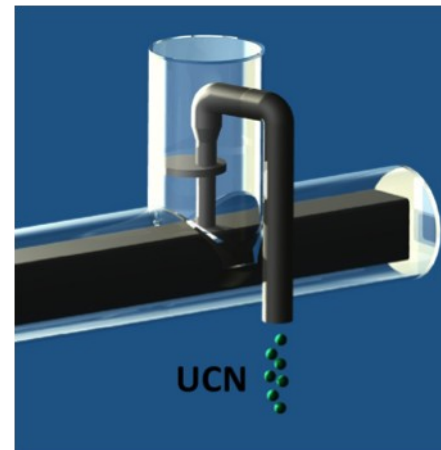
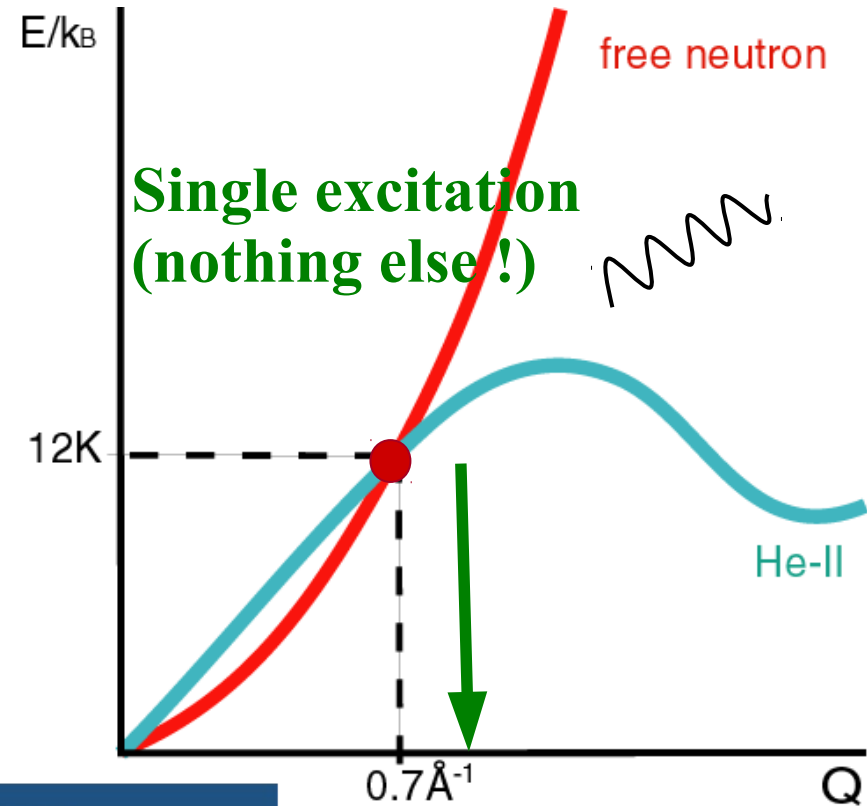
We shall use the same methodology as for water, and produce $S(\alpha, \beta)$ ACE files for MCNP.

liq-⁴He: production of UCNs

Team : O. Zimmer, J. Bossy, T. Soldner, J. Ollivier, et al.

Idea from Pendlebury and Golub in 1970's, validated in 2002.

- Crosses at $\lambda=8.9 \text{ \AA}$ for free (cold) n. Neutron loses all energy by phonon emission \rightarrow UCN.
- **Reverse suppressed** by Boltzmann factor, l-⁴He is at 0.5K, no 12K phonons.
- Extraction using open UCN converter demonstrated Y. Masuda et al., PRL (2002) 284801-1.
- Use large volume for maximum production and storage.



McStas

- We assemble a ⁴He cylinder $\phi=6.6$ cm 20 cm long (McStas).
- Send cold neutrons (velocity selector $d\lambda/\lambda=+/-8$ or 14%)
- Record escaping neutrons with $\omega < 250$ neV
- Model assembly : 1 hour (1 man) ; Computing time 1h (8 cores)

Liquid 4He

$$\sigma_{\text{abs}} = 0.00747 \text{ b (good)}$$

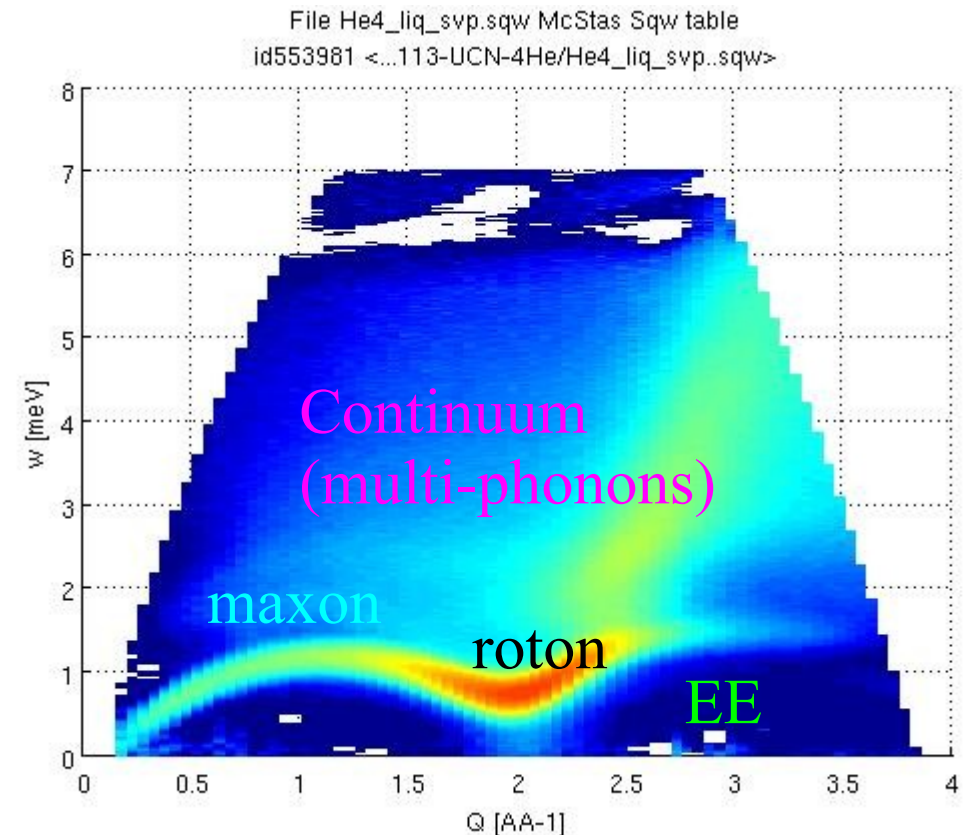
$$\sigma_{\text{coh}} = 1.34 \text{ b}$$

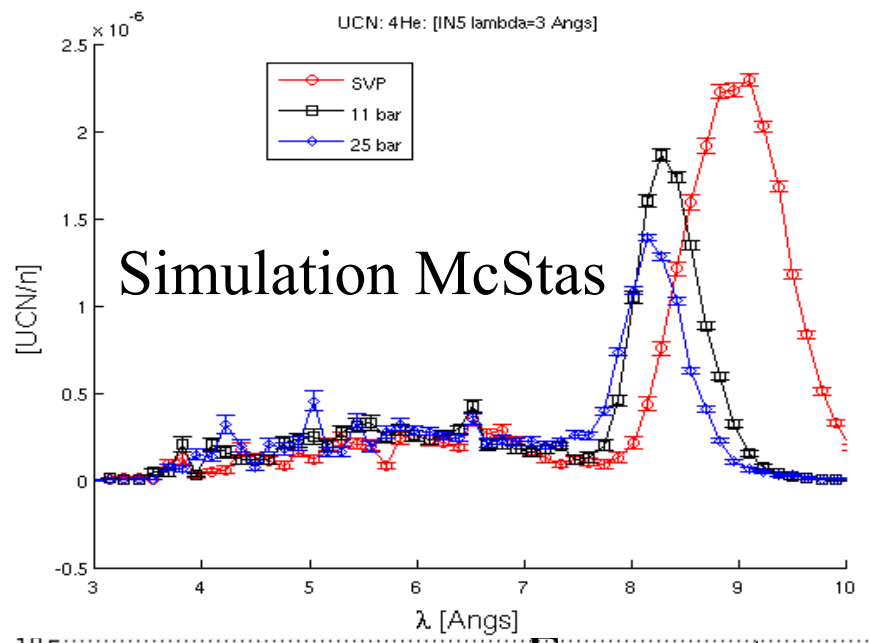
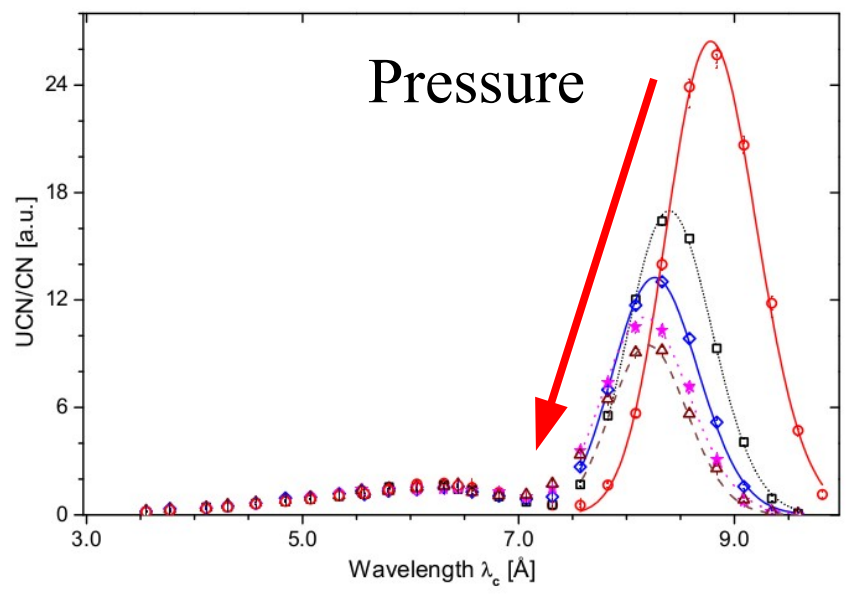
$$\sigma_{\text{inc}} = 0 \text{ b (perfect)}$$

Excellent for long storage times and large volumes.

$S(q, \omega)$ data from J. Bossy and J. Ollivier. 2.5 hours@IN5. $\lambda=3 \text{ \AA}$

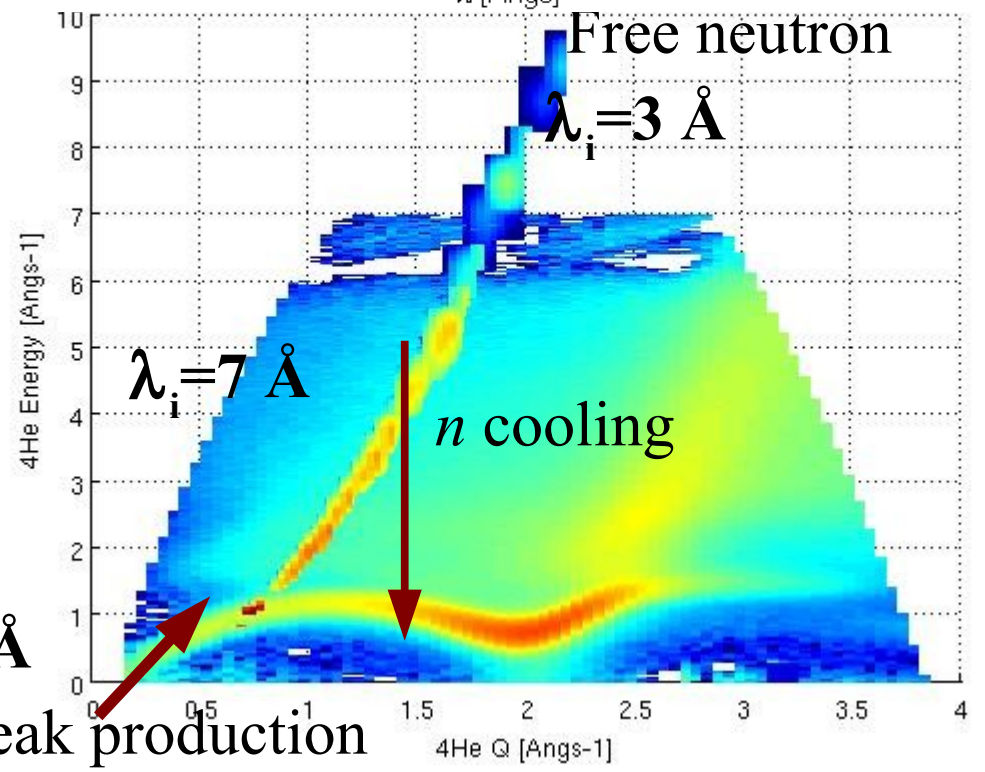
Also have IN5 $\lambda=5 \text{ \AA}$ + IN6 $\lambda=4 \text{ \AA}$





Measurement H113@ILL vs pressure

Increasing the density with pressure does not bring more neutrons as this is compensated by a smaller total cross section from steeper phonons.



McStas

- We assemble a ⁴He cylinder $\phi=6.6$ cm 20 cm long (McStas).
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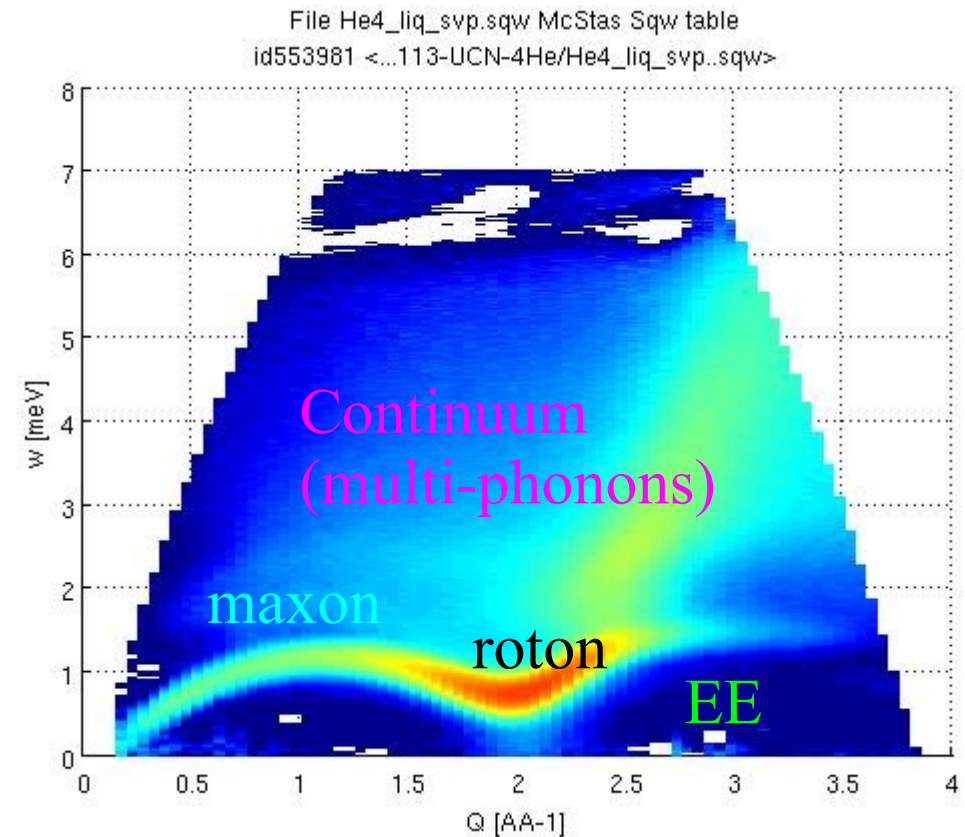
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Excellent for long storage times and large volumes.

$S(q, \omega)$ data from J. Bossy and J. Ollivier. 2.5 hours@IN5. $\lambda=3 \text{ \AA}$

Also have IN5 $\lambda=5 \text{ \AA}$ + IN6 $\lambda=4 \text{ \AA}$



McStas

- We assemble a ⁴He cylinder $\phi=6.6$ cm 20 cm long (McStas).
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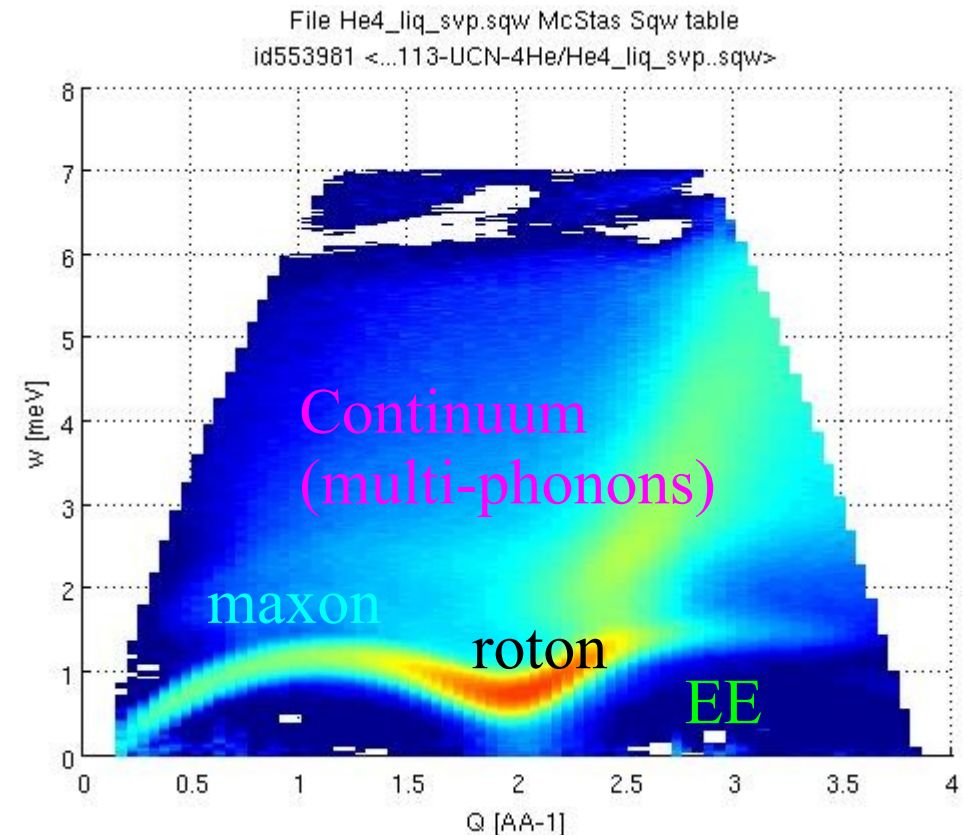
$$\sigma_{\text{coh}} = 1.34 \text{ b}$$

$$\sigma_{\text{inc}} = 0 \text{ b (perfect)}$$

Excellent for long storage times and large volumes.

$S(q, \omega)$ data from J. Bossy and J. Ollivier. 2.5 hours@IN5. $\lambda=3$ Å

Also have IN5 $\lambda=5$ Å + IN6 $\lambda=4$ Å



And that will make a day (for today at least).

